Rotation Invariant Minkowski Classes of Convex Bodies

Rolf Schneider and Franz E. Schuster

Abstract

A Minkowski class is a closed subset of the space of convex bodies in Euclidean space \mathbb{R}^n which is closed under Minkowski addition and non-negative dilatations. A convex body in \mathbb{R}^n is universal if the expansion of its support function in spherical harmonics contains non-zero harmonics of all orders. If K is universal, then a dense class of convex bodies M has the following property. There exist convex bodies T_1, T_2 such that $M + T_1 = T_2$, and T_1, T_2 belong to the rotation invariant Minkowski class generated by K. We show that every convex body K which is not centrally symmetric has a linear image, arbitrarily close to K, which is universal. A modified version of the result holds for centrally symmetric convex bodies. In this way, we strengthen a result of S. Alesker, and at the same time give a more elementary proof.

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1 Introduction and Main Results

Let \mathcal{K}^n denote the space of convex bodies (non-empty, compact, convex sets) in *n*-dimensional Euclidean space \mathbb{R}^n $(n \geq 2)$. The basic algebraic structures on \mathcal{K}^n are Minkowski addition and dilatation, defined by

$$K + L := \{x + y : x \in K, y \in L\} \quad \text{and} \quad \lambda K := \{\lambda x : x \in K\},\$$

respectively, for $K, L \in \mathcal{K}^n$ and $\lambda \geq 0$. By a *Minkowski class* in \mathbb{R}^n we understand (slightly modifying the definition of an *M*-class given in [12, p. 164]) a subset of \mathcal{K}^n which is closed in the Hausdorff metric and closed under Minkowski addition and dilatation.

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If G is a group of transformations of \mathbb{R}^n , the Minkowski class \mathcal{M} is called G-invariant if $K \in \mathcal{M}$ implies $gK \in \mathcal{M}$ for all $g \in G$. The smallest G-invariant Minkowski class containing a given convex body $K \in \mathcal{K}^n$ is said to be the G-invariant Minkowski class generated by K. It consists of all convex bodies which can be approximated by bodies of the form $\lambda_1 g_1 K + \ldots + \lambda_k g_k K$ with $k \in \mathbb{N}, \lambda_1, \ldots, \lambda_k \geq 0$, and $g_1, \ldots, g_k \in G$.

The elements of a Minkowski class \mathcal{M} will be called \mathcal{M} -bodies. Further, the convex body K is a generalized \mathcal{M} -body if there exist \mathcal{M} -bodies T_1, T_2 such that $K + T_1 = T_2$.

Let \mathcal{M} be the *G*-invariant Minkowski class generated by a convex body K. Then the convex body M is a generalized \mathcal{M} -body if and only if its support function can be approximated, uniformly on the unit sphere, by functions from the vector space spanned by the support functions of the *G*-images of K.

We recall a classical example. A convex body in \mathbb{R}^n $(n \ge 2)$ is a *zonoid* if it can be approximated by Minkowski sums of finitely many closed line segments. Thus, the set \mathcal{Z}^n of zonoids is the rigid motion invariant Minkowski class generated by a (nondegenerate) segment. Since the affine image of a segment is a segment, \mathcal{Z}^n is also the affine invariant Minkowski class generated by a segment.

Every zonoid belongs to the subset $\mathcal{K}_c^n \subset \mathcal{K}^n$ of convex bodies which have a centre of symmetry; such bodies are called *symmetric* in the following, and *origin symmetric* if 0 is the centre of symmetry. It is easy to see that $\mathcal{Z}^2 = \mathcal{K}_c^2$, but for $n \geq 3$ the set \mathcal{Z}^n is nowhere dense in \mathcal{K}_c^n . A convex body $K \subset \mathbb{R}^n$ is called a *generalized zonoid* if there exist zonoids Z_1, Z_2 such that $K + Z_1 = Z_2$. The set of generalized zonoids turns out to be dense in \mathcal{K}_c^n , see [12, Corollary 3.5.6]. Generalized zonoids played a critical role in the first author's solution [10] of the Shephard problem and in Klain's classification of translation invariant, even and simple valuations, see [7], [8]. More information on zonoids and generalized zonoids is found in the survey articles [14], [5] and in Section 3.5 of the book [12].

In the following, we want to replace the segment, which is used in the definition of zonoids, by other convex bodies. The non-denseness of zonoids mentioned above extends as follows. For $n \geq 3$, the affine invariant Minkowski class generated by a convex body (a symmetric convex body) is nowhere dense in \mathcal{K}^n (nowhere dense in \mathcal{K}^n_c). This follows from [12, Theorem 3.3.3].

Our main issue here is the question analogous to the denseness of generalized zonoids. Now we have to distinguish between convex bodies with or without a centre of symmetry. Let $\mathcal{K}_o^n \subset \mathcal{K}_c^n$ be the subset of origin symmetric convex bodies. In the following, a convex body is called *non-symmetric* if it does not have a centre of symmetry, and *non-trivial* if it has more than one point. Alesker [2] has proved the following theorem, in a different but equivalent formulation.

Theorem (Alesker). (a) If \mathcal{M} is the SL(n)-invariant Minkowski class generated by a non-symmetric convex body, then the set of generalized \mathcal{M} -bodies is dense in \mathcal{K}^n .

(b) Let \mathcal{M} be the SL(n)-invariant Minkowski class generated by a non-trivial symmetric convex body K with centre different from 0 (with centre 0). Then the set of generalized \mathcal{M} -bodies is dense in \mathcal{K}^n_c (dense in \mathcal{K}^n_o). Part (b) of this theorem extends the statement about generalized zonoids recalled above (and can be deduced from it, see [2, Remark 9]).

Note that in Alesker's result in effect the general linear group GL(n) is applied to the convex body K, since Minkowski classes are dilatation invariant. In contrast to the case of segments, for a general convex body K the affine invariant Minkowski class does not coincide with the rigid motion invariant Minkowski class generated by K. However, part (a) of Alesker's theorem should be compared to an immediate consequence of a result proved and used in [13]:

Theorem. Let $T \subset \mathbb{R}^n$ be a triangle. Then there exists an affine map A such that for the rotation invariant Minkowski class \mathcal{M} generated by AT, the set of generalized \mathcal{M} -bodies is dense in \mathcal{K}^n .

Hence, for an arbitrary triangle T, the SL(n)-invariant Minkowski class generated by T in part (a) of Alesker's Theorem may be replaced by the rotation invariant Minkowski class generated by AT, for one suitably chosen affine map A. Alesker [2], p. 58, remarks that it is not clear whether this result can be obtained by his method.

Our aim in the following is to strengthen Alesker's theorem in a way suggested by the latter theorem. Instead of applying all linear transformations to a given convex body, it is sufficient to perturb it only a little by an appropriate linear map and then to apply only rotations and dilatations.

Theorem 1. (a) Let $K \in \mathcal{K}^n$ be a non-symmetric convex body. Then there exists a linear map A, arbitrarily close to the identity, such that for the rotation invariant Minkowski class \mathcal{M} generated by AK, the set of generalized \mathcal{M} -bodies is dense in \mathcal{K}^n .

(b) Let $K \in \mathcal{K}^n$ be a non-trivial symmetric convex body, with centre different from 0 (with centre 0). Then there exists a linear map A, arbitrarily close to the identity, such that for the rotation invariant Minkowski class \mathcal{M} generated by AK, the set of generalized \mathcal{M} -bodies is dense in \mathcal{K}^n_c (dense in \mathcal{K}^n_o).

Clearly, the perturbation by the linear transformation A is necessary in general, as shown by the case of a ball in the symmetric case, and by a body of constant width in the non-symmetric case.

Besides strengthening Alesker's theorem, our second aim was to give an easier proof for it. Whereas [2] employs deep results on representations of the general linear group (proving an irreducibility theorem which is analogous to Alesker's [1] irreducibility theorem used in the theory of valuations), our proof uses spherical harmonics and is comparatively elementary and self contained.

The basic notion in our method of proof is that of universal convex bodies. A convex body K in \mathbb{R}^n is called *universal* (*centrally universal*) if the expansion of its support function in spherical harmonics contains non-zero harmonics of all orders (of all even orders). Universal convex bodies were introduced and applied by the first author in [11, p. 70]. The following theorem shows why they are crucial for our result.

Theorem 2. Let $K \in \mathcal{K}^n$ be a convex body, and let \mathcal{M} be the rotation invariant Minkowski class generated by K.

(a) The set of generalized \mathcal{M} -bodies is dense in \mathcal{K}^n if and only if K is universal.

(b) Let K be symmetric. If K has centre different from 0 (centre 0), then the set of generalized \mathcal{M} -bodies is dense in \mathcal{K}_c^n (dense in \mathcal{K}_o^n) if and only if K is centrally universal.

The only explicit convex bodies which are known to be universal are the triangles with Steiner point not at the origin and with the property that at least one of their angles is an irrational multiple of π , see [11], p. 71.

We consider the next theorem as our main result. In view of Theorem 2, it implies Theorem 1.

Theorem 3. (a) Let $K \in \mathcal{K}^n$ be a non-symmetric convex body. Then there exists a linear transformation A, arbitrarily close to the identity, such that AK is universal.

(b) Let $K \in \mathcal{K}^n$ be a non-trivial convex body. Then there exists a linear transformation A, arbitrarily close to the identity, such that AK is centrally universal.

Remark. Our proof of Theorem 3, and thus of Theorem 1, shows that, in fact, in every neighbourhood of the identity in GL(n), almost all (in the sense of measure) linear maps have the required property.

We will prove Theorem 2 in Section 2, part (b) of Theorem 3 in Section 3, and finish the proof of Theorem 3 in Section 4.

2 Universal Convex Bodies

In this section, we collect a few facts about spherical harmonics and convex bodies. An introduction to spherical harmonics and their use in convexity is found in the book of Groemer [6]; see also the short appendix of [12].

By S^{n-1} we denote the unit sphere of \mathbb{R}^n and by σ the spherical Lebesgue measure on S^{n-1} . A spherical harmonic of dimension n and order m is the restriction to S^{n-1} of a harmonic polynomial of degree m on \mathbb{R}^n . For $m \in \mathbb{N}_0$, we denote by \mathcal{H}_m^n the real vector space of spherical harmonics of dimension n and order m. \mathcal{H}^n will denote the space of all finite sums of spherical harmonics of dimension n.

 \mathcal{H}_m^n is a finite-dimensional subspace of $C(S^{n-1})$, the vector space of real continuous functions on S^{n-1} ; let N(n, m) denote its dimension. With respect to the scalar product defined by

$$(f,g) := \int_{S^{n-1}} fg \,\mathrm{d}\sigma, \qquad f,g \in C(S^{n-1}),$$

spherical harmonics of different orders are orthogonal. By

$$\pi_m: C(S^{n-1}) \to \mathcal{H}^n_m$$

we denote the orthogonal projection to \mathcal{H}_m^n . In each space \mathcal{H}_m^n we choose an orthonormal basis $\{Y_{m1}, \ldots, Y_{mN(n,m)}\}$, which will be kept fixed in the following; then

$$\pi_m f = \sum_{j=1}^{N(n,m)} (f, Y_{mj}) Y_{mj} \quad \text{for } f \in C(S^{n-1}).$$
(1)

One also writes

$$f \sim \sum_{m=0}^{\infty} \pi_m f$$

and calls this the *condensed harmonic expansion* of f (Groemer [6], p. 72). The series converges to f in the L_2 -norm.

The rotation group SO(n) acts on $C(S^{n-1})$ by means of $(\vartheta f)(u) = f(\vartheta^{-1}u), u \in S^{n-1}$. The space \mathcal{H}_m^n is invariant under rotations. Thus, for any rotation $\vartheta \in SO(n)$, we have

$$\vartheta Y_{mj}(u) = \sum_{i=1}^{N(n,m)} t^m_{ij}(\vartheta) Y_{mi}(u), \qquad u \in S^{n-1},$$
(2)

with real coefficients $t_{ij}^m(\vartheta)$. Let ν denote the normalized Haar measure on the compact group SO(n). The following formula was proved in [13, Lemma 3]. If $f \in C(S^{n-1})$, then

$$\int_{SO(n)} \vartheta f(u) t_{ij}^m(\vartheta) \, \mathrm{d}\nu(\vartheta) = N(n,m)^{-1}(f,Y_{mj}) Y_{mi}(u) \tag{3}$$

for $u \in S^{n-1}, n \in \mathbb{N}_0$, and i, j = 1, ..., N(n, m).

A convex body $K \in \mathcal{K}^n$ is determined by its support function $h(K, \cdot)$, defined on \mathbb{R}^n by $h(K, x) = \max\{\langle x, y \rangle : y \in K\}$. Since $h(K, \cdot)$ is positively homogeneous of degree one, it is determined by its restriction to the sphere S^{n-1} , which we denote by h_K . If $K, L \in \mathcal{K}^n$, then $h(K + L, \cdot) = h(K, \cdot) + h(L, \cdot)$. Moreover, convergence in the Hausdorff metric on \mathcal{K}^n is equivalent to uniform convergence of support functions on S^{n-1} .

The functions $\pi_m h_K$, $m \in \mathbb{N}$, determine the convex body K uniquely. In particular, a convex body K is symmetric if and only if $\pi_m h_K = 0$ for all odd numbers $m \neq 1$, and K is one-pointed if and only if $h_K \in \mathcal{H}_1^n$. It follows that for $m \neq 1$, the projection $\pi_m h_K$ is invariant under translations of K. We also note two special cases:

$$(\pi_0 h_K)(u) = \frac{1}{2}b(K)$$
 and $(\pi_1 h_K)(u) = \langle s(K), u \rangle = h_{\{s(K)\}}(u)$

for $u \in S^{n-1}$ and $K \in \mathcal{K}^n$. Here, $b : \mathcal{K}^n \to \mathbb{R}$ is the mean width of the convex body K, defined by

$$b(K) = \frac{2}{\omega_n} \int_{S^{n-1}} h_K \,\mathrm{d}\sigma,$$

where $\omega_n = \sigma(S^{n-1})$. The rigid motion equivariant map $s : \mathcal{K}^n \to \mathbb{R}^n$ is the Steiner point map, defined by

$$s(K) = \frac{n}{\omega_n} \int_{S^{n-1}} h_K(u) u \, \mathrm{d}\sigma(u).$$

It follows that $\pi_0 h_K = 0$ if and only if K contains only one point, and $\pi_1 h_K = 0$ if and only if s(K) = 0.

We can now give a precise definition of universal convex bodies.

Definition. The convex body $K \in \mathcal{K}^n$ is called *universal* if $\pi_m h_K \neq 0$ for all $m \in \mathbb{N}_0$. The body $K \in \mathcal{K}^n$ is *centrally universal* if $\pi_m h_K \neq 0$ holds for all even numbers $m \in \mathbb{N}_0$.

Since the space \mathcal{H}_m^n and the scalar product on $C(S^{n-1})$ are invariant under rotations, we have $\pi_m h_{\lambda\vartheta K} = \lambda \vartheta(\pi_m h_K)$ for every $\vartheta \in SO(n)$ and every $\lambda \geq 0$. Therefore, the property of being universal (centrally universal) is invariant under rotations and dilatations.

Let $K \in \mathcal{K}^n$ be a k-dimensional convex body, $k \geq 2$. Because of the rotation invariance just mentioned and the translation equivariance of the Steiner point map, it is no loss of generality to assume that $K \subset \mathbb{R}^k \subset \mathbb{R}^n$. If K is universal in the sense of the preceding definition, we say that K is *universal in* \mathbb{R}^n . Alternatively, we can consider K as a subset of the Euclidean space \mathbb{R}^k . If it is universal there, that is, in the sense of the definition with n = k, we say that K is *universal in* \mathbb{R}^k . The following was proved in [11, §5].

Lemma 1. If K is universal in \mathbb{R}^k , then K is universal in \mathbb{R}^n .

Before we turn to the proof of Theorem 2, we recall that the set of convex bodies $L \in \mathcal{K}^n$ with $h_L \in \mathcal{H}^n$ is dense in \mathcal{K}^n , see [12, p. 160]. Similarly, the set of all $L \in \mathcal{K}^n_c$ with $h_L \in \mathcal{H}^n$ is dense in \mathcal{K}^n_c , and the set of all $L \in \mathcal{K}^n_o$ with $h_L \in \mathcal{H}^n$ is dense in \mathcal{K}^n_c .

Proof of Theorem 2. (a) Suppose K is not universal, i.e., $\pi_m h_K = 0$ for some m. If m = 0, then K is one-pointed, hence we can assume that $m \ge 1$. Let Y_m be a non-zero spherical harmonic of order m. There is a constant c > 0 such that $c + Y_m = h_M$ for some convex body M, see [12, Lemma 1.7.9]. Every function f in the closure of the vector space spanned by the functions $h_{\vartheta K}$, $\vartheta \in SO(n)$, satisfies $\pi_m f = 0$. Since this does not hold for h_M , the set of generalized \mathcal{M} -bodies is not dense in \mathcal{K}^n .

Now let K be universal. By the preceding remark, it is enough to show that for every convex body L with $h_L \in \mathcal{H}^n$ there are \mathcal{M} -bodies T_1, T_2 such that $L + T_1 = T_2$. This was proved in [13] for the case where K is a triangle with at least one of its angles an irrational multiple of π . For an arbitrary universal convex body K, the proof is almost verbally the same. We sketch the argument, for the reader's convenience and since we have to explain the modifications in part (b). Let

$$h_L = \sum_{m=0}^k \sum_{j=1}^{N(n,m)} a_{mj} Y_{mj}$$
(4)

and define $c_{mj} = (h_K, Y_{mj})$. Since K is universal, for each $m \in \mathbb{N}_0$ there is an index j(m) such that $c_{mj(m)} \neq 0$. With the functions t_{ij}^m from (2), we define

$$g := N(n,m) \sum_{m=0}^{k} \sum_{i,j=1}^{N(n,m)} b_{ij}^{m} t_{ij}^{m},$$
(5)

where $b_{ij}^m = a_{mi} c_{mj(m)}^{-1}$ for j = j(m) and 0 otherwise. By (3), we have

$$\int_{SO(n)} h_{\vartheta K}(u)g(\vartheta) \, d\nu(\vartheta) = N(n,m) \sum_{m=0}^{k} \sum_{i,j=1}^{N(n,m)} \int_{SO(n)} h_{\vartheta K}(u) b_{ij}^{m} t_{ij}^{m}(\vartheta) \, d\nu(\vartheta)$$
$$= \sum_{m=0}^{k} \sum_{i,j=1}^{N(n,m)} b_{ij}^{m}(h_{K}, Y_{mj}) Y_{mi}(u)$$
$$= \sum_{m=0}^{k} \sum_{i=1}^{N(n,m)} a_{mi} Y_{mi}(u) = h_{L}(u).$$

Splitting g into positive and negative parts, we obtain that L is a generalized \mathcal{M} -body.

(b) If K is not centrally universal, then $\pi_m h_K$ for some even number m. With M constructed as in part (a), we have $M \in \mathcal{K}_o^n$, but M cannot be approximated by bodies from \mathcal{M} , since all bodies $L \in \mathcal{M}$ satisfy $\pi_m h_L = 0$.

Suppose that K is centrally universal. First let K have its centre different from 0. Then $\pi_1 h_K \neq 0$. Let $L \in \mathcal{K}_c^n$ and $h_L \in \mathcal{H}^n$. Let h_L be represented by (4); then $a_{mj} = 0$ for all odd $m \neq 1$. For even numbers m and for m = 1, we can define c_{mj} and b_{ij}^m as in part (a), as well as the function g, where now $b_{ij}^m := 0$ for odd $m \neq 1$.

If K has centre 0, then $\pi_1 h_K = 0$ and we choose also $b_{ij}^m := 0$ for m = 1. The proof can now be completed as in part (a).

3 Linear transformations

In this section, we investigate the behaviour of $\pi_m h_{AK}$ under linear transformations A.

Lemma 2. Let $f, g : \mathbb{R}^n \setminus \{0\} \to \mathbb{R}$ be continuous functions, let f be positively homogeneous of degree 1 and g positively homogeneous of degree -(n+1). Then, for every $A \in GL(n)$,

$$\int_{S^{n-1}} f(Av)g(v) \, \mathrm{d}\sigma(v) = \frac{1}{|\det A|} \int_{S^{n-1}} f(v)g(A^{-1}v) \, \mathrm{d}\sigma(v).$$

Proof. We can take advantage of the known transformation behaviour of the dual mixed volume $\tilde{V}_{-1}(K, L)$. It is defined for two star bodies $K, L \subset \mathbb{R}^n$ (i.e., compact sets, starshaped with respect to the origin and with continuous radial functions) by

$$\tilde{V}_{-1}(K,L) = \frac{1}{n} \int_{S^{n-1}} \rho(K,u)^{n+1} \rho(L,u)^{-1} \,\mathrm{d}\sigma(u),$$

where $\rho(K, u) = \max\{\lambda \ge 0 : \lambda u \in K\}$ denotes the radial function of K. Decomposing f into its positive and negative part and adding, say, the function $x \mapsto ||x||$ to both components, we can write $f = f^+ - f^-$, where the functions f^+, f^- are positive,

continuous, and positively homogeneous of degree 1, on $\mathbb{R}^n \setminus \{0\}$. Let L_+, L_- be the star bodies with radial functions $\rho(L_+, \cdot) = (f^+)^{-1}$ and $\rho(L_-, \cdot) = (f^-)^{-1}$, then

$$f = \rho(L_+, \cdot)^{-1} - \rho(L_-, \cdot)^{-1}.$$

Similarly, there are star bodies K_+, K_- with

$$g = \rho(K_+, \cdot)^{n+1} - \rho(K_-, \cdot)^{n+1}.$$

The assertion of the lemma now follows from the fact that $\rho(K, Av) = \rho(A^{-1}K, v)$ and that

$$V_{-1}(AK, AL) = |\det A| V_{-1}(K, L),$$

which is proved in Lutwak [9, Lemma (7.9)].

Lemma 2 is applied in the following argument. For $m \in \mathbb{N}_0$ and $j \in \{1, \ldots, N(n, m)\}$ we define

$$\check{Y}_{mj}(x) := \frac{1}{\|x\|^{n+1}} Y_{mj}\left(\frac{x}{\|x\|}\right) \quad \text{for } x \in \mathbb{R}^n \setminus \{0\}.$$

Let $K \in \mathcal{K}^n$ be fixed, and let $A \in GL(n)$. Since $h(AK, v) = h(K, A^T v)$, Lemma 2 yields

$$(h_{AK}, Y_{mj}) = \frac{1}{|\det A|} \int_{S^{n-1}} h(K, v) \check{Y}_{mj}(A^{-T}v) \, \mathrm{d}\sigma(v).$$

Since the spherical harmonic Y_{mj} is the restriction of a polynomial on \mathbb{R}^n to the sphere S^{n-1} , the function

$$A \mapsto \frac{1}{|\det A|} \check{Y}_{mj}(A^{-T}v)$$

is real analytic on the connected component of the identity. (We identify GL(n), via matrix description, with an open subset of \mathbb{R}^{n^2} .) The convergence of its power series is uniform on every compact subset and uniform in v. Using the compactness of S^{n-1} , we see that the function defined by

$$A \mapsto (h_{AK}, Y_{mj}),$$

(for given K) is real analytic. We will mostly apply this with linear maps $A(\lambda)$ which, with respect to the standard orthonormal basis of \mathbb{R}^n , have diagonal matrices diag $(1, \lambda, \ldots, \lambda)$ or (in Section 4) diag $(1, 1, \lambda, \ldots, \lambda)$, with $\lambda \in I$. Here I is any open interval (0, a) with a > 1. Then the function defined by

$$f_{mj}(\lambda) := (h_{A(\lambda)K}, Y_{mj}), \qquad \lambda \in I,$$
(6)

is real analytic on I.

We are now in a position to prove part (b) of Theorem 3. This case exhibits already the basic idea for the proof of the general result.

Proof of Theorem 3 (b). Let $K \in \mathcal{K}_c^n$ be a non-trivial convex body. For $\lambda \in I$, let $A(\lambda) \in GL(n)$ be defined by $A(\lambda) : (x_1, \ldots, x_n) \mapsto (x_1, \lambda x_2, \ldots, \lambda x_n)$ (where x_1, \ldots, x_n

are coordinates with respect to the standard orthonormal basis of \mathbb{R}^n). We may assume that the orthogonal projection of K to the first coordinate axis is a segment S of positive length. We have $\lim_{\lambda\to 0} A(\lambda)K = S$ in the Hausdorff metric and hence

$$\lim_{\lambda \to 0} (h_{A(\lambda)K}, Y_{mj}) = (h_S, Y_{mj})$$

for all m, j. Let m be even. It is well known that, for all $Y_m \in \mathcal{H}_m^n$, $u \in S^{n-1}$,

$$\int_{S^{n-1}} |\langle u, v \rangle| Y_m(v) \, \mathrm{d}\sigma(v) = a_m Y_m(u)$$

with $a_m \neq 0$ (see, e.g., [12, p. 185]). Since $h_S = |\langle x, \cdot \rangle| + \langle y, \cdot \rangle$ for suitable $x, y \in \mathbb{R}^n$ it follows that S is centrally universal. Therefore, for each even $m \in \mathbb{N}$ there is at least one index $j(m) \in \{1, \ldots, N(n, m)\}$ for which $(h_S, Y_{mj(m)}) \neq 0$. This implies that the function $f_{mj(m)}$, defined by (6), does not vanish identically on I. Since it is real analytic, its zeros are isolated, thus there is an at most countable subset $Z_m \subset I$ such that $f_{mj(m)}(\lambda) \neq 0$ for $\lambda \in I \setminus Z_m$. For such λ , we have $\pi_m h_{A(\lambda)K} \neq 0$ (this holds trivially for m = 0 and all $\lambda \in I$, since $A(\lambda)K$ has positive mean width). If now $U \subset \mathbb{R}$ is a given neighbourhood of 1, there exists a number

$$\lambda \in U \setminus \bigcup_{m \in \mathbb{N}, \, 2|m} Z_m.$$

It satisfies $\pi_m h_{A(\lambda)K} \neq 0$ for all even m, hence $A(\lambda)K$ is centrally universal. Such a map $A(\lambda) \in GL(n)$ can be found in any prescribed neighbourhood of the identity. This completes the proof of Theorem 3 (b).

4 Proof of Theorem 3 (a)

For the proof of Theorem 3 (a), we first treat the case n = 2. Let $K \in \mathcal{K}^2$ be a non-symmetric convex body; then K has interior points.

As usual, we parameterize S^1 by an angle, writing $u_{\varphi} := (\cos \varphi, \sin \varphi)$, and by slight abuse of notation, for a function f on S^1 we write $f(u_{\varphi}) = f(\varphi)$. The space \mathcal{H}_m^2 is spanned by the functions $\cos m\varphi$ and $\sin m\varphi$, thus, in complex notation,

$$\pi_m h_K = 0 \iff \int_0^{2\pi} h(K, \varphi) e^{im\varphi} d\varphi = 0.$$

For $m \in \mathbb{N}$, we define a map $F_m(K, \cdot) : GL(2)^+ \to \mathbb{C}$ by

$$F_m(K,A) := \int_0^{2\pi} h(AK,\varphi) e^{im\varphi} d\varphi \quad \text{for } A \in GL(2)^+,$$

where $GL(2)^+$ denotes the connected component of the identity in GL(2). As explained in Section 3, the function $F_m(K, \cdot)$ (interpreted as a function on an open subset of \mathbb{R}^4) is real analytic. We make use of the following fact (for a proof, see [3, Lemma 5]). **Lemma 3.** Let $f: U \to \mathbb{R}$ be a real analytic function on an open subset of \mathbb{R}^k . Then the zero set of f has Lebesgue measure zero, unless f vanishes identically.

In view of Lemma 3, the following auxiliary result will be crucial.

Lemma 4. The relation

$$F_m(K,\cdot) \equiv 0$$

does not hold for any odd integer $m \geq 1$.

For the proof, we assume that the assertion were false. Then there exists a smallest odd integer $m \ge 1$ with $F_m(K, \cdot) \equiv 0$.

First Case: $m \ge 5$. By the choice of m, we have $F_{m-2}(K, \cdot) \ne 0$, hence there exists a map $A_0 \in GL(2)^+$ with $F_{m-2}(K, A_0) \ne 0$, equivalently $F_{m-2}(A_0K, \mathrm{Id}) \ne 0$. We still have $F_m(A_0K, \cdot) \equiv 0$. We may now replace K by A_0K and change the notation. Thus, we can assume that

$$F_m(K, \cdot) \equiv 0, \qquad F_{m-2}(K, \mathrm{Id}) \neq 0. \tag{7}$$

Second Case: m = 1 or m = 3. Since K is not symmetric, we cannot have $F_k(K, \mathrm{Id}) = 0$ for all odd integers $k \geq 3$. A fortiori, $F_k(K, \cdot) \equiv 0$ cannot hold for all odd integers $k \geq 3$. Therefore, there exists an odd integer $k \geq 3$ such that $F_{k-2}(K, \cdot) \equiv 0$, but $F_k(K, \cdot) \not\equiv 0$. As in the first case, we can replace K by A_0K , with a linear map A_0 , so that after a change of notation we have

$$F_{k-2}(K, \cdot) \equiv 0, \quad \text{but} \quad F_k(K, \text{Id}) \neq 0.$$
 (8)

For a while, both cases are now treated together. To K we will apply rotations $R(\alpha)$ and linear maps $A(\lambda)$, given by matrices

$$\begin{pmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{pmatrix}$$
 and $\begin{pmatrix} 1 & 0 \\ 0 & \lambda \end{pmatrix}$

respectively, with α in an open neighbourhood U of 0 and $\lambda \in I$. We have

$$h(A(\lambda)K, u_{\varphi}) = h(K, A(\lambda)u_{\varphi}) = ||A(\lambda)u_{\varphi}||h(K, u_{\psi})$$

with

$$u_{\psi} := \frac{A(\lambda)u_{\varphi}}{\|A(\lambda)u_{\varphi}\|} = \frac{(\cos\varphi,\lambda\sin\varphi)}{\sqrt{\cos^{2}\varphi + \lambda^{2}\sin^{2}\varphi}}.$$

Then

$$u_{\varphi} = \frac{(\lambda \cos \psi, \sin \psi)}{\sqrt{\lambda^2 \cos^2 \psi + \sin^2 \psi}}$$

thus

$$e^{im\varphi} = \frac{(\lambda\cos\psi + i\sin\psi)^m}{(\lambda^2\cos^2\psi + \sin^2\psi)^{\frac{m}{2}}}$$

and

$$\frac{\mathrm{d}\varphi}{\mathrm{d}\psi} = \frac{\lambda}{\lambda^2 \cos^2 \psi + \sin^2 \psi}.$$

Substituting φ by ψ in the integral defining $F_m(K, A)$, we get

$$F_m(K, A(\lambda)) = \lambda^2 \int_0^{2\pi} h(K, \psi) \frac{(\lambda \cos \psi + i \sin \psi)^m}{(\lambda^2 \cos^2 \psi + \sin^2 \psi)^{\frac{m+3}{2}}} d\psi.$$

By (7), this integral vanishes for all $\lambda \in I$. Therefore, also its derivatives with respect to λ vanish. For

$$f(\lambda) = \frac{(\lambda \cos \psi + i \sin \psi)^m}{(\lambda^2 \cos^2 \psi + \sin^2 \psi)^{\frac{m+3}{2}}}$$

we obtain

$$0 = -f'(1) = \frac{3}{2}e^{im\psi} + \frac{3-m}{4}e^{i(m-2)\psi} + \frac{3+m}{4}e^{i(m+2)\psi}.$$

Since $F_m(K, \mathrm{Id}) = 0$, this yields

$$\int_{0}^{2\pi} h(K,\psi) [(3-m) e^{i(m-2)\psi} + (3+m) e^{i(m+2)\psi}] d\psi = 0.$$
(9)

According to (7), we also have $F_m(R(\alpha)K, \cdot) \equiv 0$ for each angle $\alpha \in U$, hence (9) holds if K is replaced by $R(\alpha)K$. Since $h(R(\alpha)K, \psi) = h(K, \psi - \alpha)$, we see after a substitution that (9) holds with ψ in the exponentials replaced by $\psi + \alpha$. Since the functions $e^{i(m-2)\alpha}$ and $e^{i(m+2)\alpha}$ are linearly independent on U, we deduce that

$$F_{m-2}(K, \mathrm{Id}) = 0 \quad \text{if } m \neq 3, \qquad F_{m+2}(K, \mathrm{Id}) = 0.$$
 (10)

Now we distinguish between the two cases considered above. If $m \geq 5$, then the first relation of (10) yields $F_{m-2}(K, \mathrm{Id}) = 0$, which contradicts the second relation of (7). If m = 1 or m = 3, then we note that the first relation of (8) gives $F_{k-2}(K, \cdot) \equiv 0$. Therefore, the second relation of (10) holds also with m replaced by k - 2, but this contradicts the second relation of (8). This completes the proof of Lemma 4.

Let $m \in \mathbb{N}_0$ be an integer. If m is odd, it follows from Lemma 4 that the real analytic function $F_m(K, \cdot)$ does not vanish identically. If m is even, the same result follows as in the proof of Theorem 3 (b). Hence, the set of zeros of $F_m(K, \cdot)$ has Lebesgue measure zero, by Lemma 3. Therefore, in any given neighbourhood of the identity in GL(2), we can find a linear map A with $F_m(K, A) \neq 0$ for all $m \in \mathbb{N}_0$. The convex body AK is universal. This completes the proof of Theorem 3 for n = 2 and non-symmetric convex bodies.

Finally, we assume that $n \geq 3$ and that $K \in \mathcal{K}^n$ is a non-symmetric convex body. Then K has dimension at least two. There exists (see, e.g., Gardner [4, Corollary 3.1.5]) a two-dimensional subspace, without loss of generality the space $\mathbb{R}^2 \subset \mathbb{R}^n$, such that the orthogonal projection K' of K to \mathbb{R}^2 is non-symmetric. In a given neighbourhood of the identity of GL(n) we can find an affine transformation B which maps \mathbb{R}^2 into itself, leaves the orthogonal complement of \mathbb{R}^2 in \mathbb{R}^n pointwise fixed, and is such that BK' is universal in \mathbb{R}^2 . By Lemma 1, BK' is universal in \mathbb{R}^n . Moreover, BK' is the image of BK under the orthogonal projection to \mathbb{R}^2 . Assuming that \mathbb{R}^2 is spanned by the first two vectors of the standard orthonormal basis of \mathbb{R}^n , we define linear maps $A(\lambda)$ by $A(\lambda) : (x_1, \ldots, x_n) \mapsto (x_1, x_2, \lambda x_3, \ldots, \lambda x_n), \lambda \in I$. As in the proof of Theorem 3 (b), we have

$$\lim_{\lambda \to 0} (h_{A(\lambda)BK}, Y_{mj}) = (h_{BK'}, Y_{mj}).$$

Since BK' is universal in \mathbb{R}^n , for each m there exists an index j(m) with $(h_{BK'}, Y_{mj(m)}) \neq 0$. Thus, the function $\lambda \mapsto (h_{A(\lambda)BK}, Y_{mj(m)}), \lambda \in I$, does not vanish identically. The proof can now be completed as it was done in Section 3 for centrally symmetric bodies.

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Professor R. Schneider, Mathematisches Institut, Albert-Ludwigs-Universität, Eckerstraße 1, D-79104 Freiburg i. Br., Germany. E-mail: rolf.schneider@math.uni-freiburg.de

Dr. F. E. Schuster, Institut für Diskrete Mathematik und Geometrie, Technische Universität Wien, Wiedner Hauptstraße 8/104, A-1040 Wien, Austria. E-mail: fschuster@osiris.tuwien.ac.at